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Experimental study of ion heating and acceleration during magnetic reconnection

S. C. Hsu,* T. A. Carter, G. Fiksel† H. Ji, R. M. Kulsrud, and M. Yamada

Princeton Plasma Physics Laboratory, P. O. Box 451, Princeton, NJ 08543

Abstract

Ion heating and acceleration has been studied in the well-characterized reconnection layer of the Magnetic Reconnection Experiment [M. Yamada et al., Phys. Plasmas 4, 1936 (1997)]. Ion temperature in the layer rises substantially during null-helicity reconnection in which reconnecting field lines are anti-parallel. The plasma outflow is sub-Alfvénic due to a downstream back pressure. An ion energy balance calculation based on the data and including classical viscous heating indicates that the ions are heated largely due to non-classical mechanisms. The $T_i$ rise is much smaller during co-helicity reconnection in which field lines reconnect obliquely. This is consistent with a slower reconnection rate and a smaller resistivity enhancement over the Spitzer value. These observations indicate strongly that non-classical dissipation mechanisms can play an important role both in heating the ions and in facilitating the reconnection process.

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*Present address: Dept. of Applied Physics, California Institute of Technology, Pasadena, CA 91125; email: scotthsu@caltech.edu.

†Permanent address: Dept. of Physics, University of Wisconsin, Madison, WI 53706.
I. INTRODUCTION

Magnetic reconnection,\textsuperscript{1,2} the topological rearrangement and annihilation of magnetic flux in a highly conductive plasma, is thought to play a central role in coronal heating, magnetospheric dynamics, and “anomalous” ion heating in reversed-field pinches (RFP) and spheromaks. Although reconnection is invoked frequently as an ion heating or acceleration mechanism, a detailed understanding of the energy conversion process is still lacking, largely due to the limited accessibility of reconnection events in both naturally occurring and laboratory plasmas. Description of the process is often limited to “black box” phenomenology, in which the input parameters are, for example, implications for magnetic nulls (such as sunspot groups) or measurable magnetic fluctuations (as in RFP’s), and the output parameters are energetic particle beams or enhanced global ion heating, respectively. The details of this black box, however, are fundamental to plasma physics since reconnection is a virtually unavoidable process in all magnetized plasmas. Because local reconnection changes global magnetic field topology, overall plasma equilibrium and confinement properties can be profoundly affected. A well-known example of this is the interaction of the solar wind with the Earth’s magnetosphere, in which reconnection along the day-side magnetopause and the polar cusp regions allow solar wind plasma to penetrate into the magnetospheric cavity, occasionally leading to large-scale geomagnetic storms. Detailed understanding of the magnetic energy conversion process could also be the first step toward eventual creative application of such understanding to alternative concept fusion or the minimization and control of confinement-degrading effects in tokamak plasmas.

Recent research on the Magnetic Reconnection Experiment\textsuperscript{3} (MRX) has focused on ion heating and acceleration during magnetic reconnection.\textsuperscript{4,5} Using a novel spectroscopy probe\textsuperscript{6} developed at the Univ. of Wisconsin–Madison, local ion temperature was measured in the well-characterized reconnection layer of MRX. In the two-dimensional (2D), quasi steady-state reconnection layer of MRX, it is possible to capture cleanly one possible scenario of what reconnection does to a real plasma when magnetic flux is annihilated. Important ques-
tions addressed include: (1) whether ions are heated by reconnection, (2) whether energetic plasma flows develop, (3) what fraction of the dissipated magnetic energy is converted to ion energy, and (4) what is the nature of the energy conversion mechanism(s). It is found that ion temperature $T_i$ rises by a factor of three in the reconnection layer during null-helicity reconnection (negligible out-of-plane field) and, furthermore, that the $T_i$ rise is causally linked to the reconnection process. The well-known Alfvénic plasma flows of classical MHD reconnection models$^{7,8}$ are not observed in MRX due to a downstream back pressure. An ion energy balance calculation, based on experimental data and including classical viscous heating by the sub-Alfvénic flows, indicates that most of the ion heating occurs via non-classical mechanisms. The $T_i$ rise is much smaller during co-helicity reconnection (finite out-of-plane field), consistent with the slower reconnection rate and a lower value of resistivity enhancement over the classical Spitzer value for co-helicity compared to null-helicity reconnection.

It is possible that non-classical dissipation mechanisms present in the null-helicity case are reduced in the low-beta, co-helicity case. These experimental observations indicate strongly that non-classical dissipation mechanisms play an important role both in heating ions and in facilitating the reconnection process.

It is important to note that prior to this work, local $T_i$ measurements had not been made in conjunction with well-characterized reconnection events, although correlations between reconnection and global ion heating as well as energetic ion flows have been reported on TS-3 at the Univ. of Tokyo$^9$ and on SSX (Swarthmore Spheromak Experiment)$^{10}$ in the prior experiments, reconnection was generated by colliding two spheromaks together at a substantial fraction of the Alfvén speed. Possible strong compressional heating during the collision could complicate the interpretation of any observed ion heating and acceleration. The slow formation scheme of MRX allows a quasi steady-state reconnection layer to form, with magnetic energy dissipation being the dominant heating source. The other main distinction between the present work and prior experiments is the direct local measurement of the majority ion temperature. Prior results were based on line-of-sight averaged neutral and impurity emission$^9$ or ion flux at the vacuum wall.$^{10}$
This paper is organized as follows. Section II describes the experimental setup, plasma formation and configurations, and diagnostics. Section III presents $T_i$ measurements which show a rise in $T_i$ that is causally linked to reconnection, both temporally and spatially. In this section, it is also shown that the $T_i$ rise is much smaller for co-helicity compared to null-helicity reconnection. Section IV presents measurements of both downstream and toroidal ion flow $V_i$, showing that flows which develop in either direction are sub-Alfvénic in magnitude. Based on the experimental data, an ion energy balance is calculated in Sec. V, showing that classical viscous heating is small and that the ions are heated predominantly due to non-classical mechanisms. Section VI discusses the relationship between ion heating and resistivity enhancement and provides a discussion on the possible non-classical ion heating mechanisms. The paper concludes with a summary.

II. EXPERIMENTAL APPARATUS

Experiments were performed on MRX$^3$ which produces plasmas satisfying the MHD criteria globally ($S \gg 1$, $\rho_i \ll L$, $V_A \ll c$, where $S$ is the Lundquist number, $\rho_i$ the ion gyro-radius, $L$ the characteristic length of the reconnection layer, $V_A$ the Alfvén speed, and $c$ the speed of light). Formation and control of the reconnection layer in MRX is unique among laboratory experiments. “Flux-cores” containing toroidal field (TF) and poloidal field (PF) coils, shown in Fig. 1(a), allow the controlled formation of a toroidal, axisymmetric reconnection layer, shown schematically in Fig. 1(b). The arrows represent the motion of plasma and magnetic flux during “pull” reconnection, in which “public flux” linking both flux-cores is reconnected into “private flux” linking each individual flux-core. This sequence produces a reconnection layer elongated in $Z$ which persists for approximately ten Alfvén transit times ($\tau_A \equiv L/V_A \approx 1.5 \mu$s). When the TF coils are connected with opposite (same) polarity, the out-of-plane field $B_\theta$ is negligible (finite), resulting in null-helicity (co-helicity) reconnection. For more details on the experimental apparatus, please see Ref. 3. Previously, detailed magnetic field measurements from the reconnection layer
were reported,\textsuperscript{11} revealing Y-shaped (O-shaped) diffusion regions in the null-helicity (co-
helicity) case and a neutral sheet half-thickness $\delta \sim c/\omega_{pi} \sim \rho_i \sim 2\text{ cm}$. The reconnection
rate was found to be consistent with a generalized Sweet-Parker model\textsuperscript{12,13} in which the
effects of compressibility, downstream pressure, and non-classical resistivity are considered.
More recently, the profile and thickness of the neutral sheet in MRX was shown\textsuperscript{14} to agree
with a modified Harris sheet equilibrium.\textsuperscript{15}

In MRX, all three components of $\mathbf{B}$ are measured using magnetic probe arrays, and
electron density $n_e$ and temperature $T_e$ are measured using a triple Langmuir probe.
Other reconnection quantities are inferred from the direct measurements, including the
poloidal flux $\psi \equiv \int_0^R 2\pi R' B_Z(R') dR'$ (assuming axisymmetry), reconnection layer current
density $j_\theta \approx -(\partial B_z/\partial R)/\mu_0$, reconnection electric field $E_\theta = -(\partial \psi/\partial t)/2\pi R$, resistivity
in the reconnection layer $\eta_* \equiv E_\theta/j_\theta$, Spitzer resistivity $\eta \sim T_e^{-3/2}$, Alfvén speed
$V_A \equiv B/\sqrt{\mu_0 \rho}$ (using the reconnecting $B$ at the edge of the layer), and plasma inflow
speed $V_R \approx -(\partial \psi/\partial t)/(\partial \psi/\partial R)$ (invoking flux-freezing outside the diffusion region). The
experiments reported here use pure helium discharges in order to achieve a direct measure-
ment of $T_i$ via Doppler spectroscopy of the He II 4686 Å line. Parameters for the present
experiments are as follows: $n_e \approx 5 \times 10^{13}$ cm$^{-3}$, $T_e \approx 10$–15 eV, and $T_i \approx 5$–15 eV (all in
the reconnection layer), $B \approx 250$ G at the edge of the layer, and $S \equiv \mu_0 LV_A/\eta \approx 250$–500.
Qualitative features of the reconnection layer for the present helium discharges are similar
to those of previously reported hydrogen discharges,\textsuperscript{11,3} except $\delta_{\text{He}} \approx 2\delta_{\text{H}}$, consistent with
the previously reported $\delta \sim \rho_i$ scaling.\textsuperscript{11} Spatially, the reconnection region is located in an
area given by $R \approx 35 \rightarrow 40$ cm and $Z \approx -10 \rightarrow 10$ cm.

A major new accomplishment of this work is the local measurement of $T_i$ using the
Ion Dynamics Spectroscopy Probe\textsuperscript{6} (IDSP) developed at UW–Madison. The IDSP is an
insertable probe which collects plasma light from a localized volume. Two perpendicular
lines-of-sight can give simultaneous Doppler broadening and relative Doppler shift inform-
ation, although only one line-of-sight is used for the present experiments. Figure 1(b)
shows the placement of the IDSP in MRX (to scale); the lines-of-sight can be oriented in
an $R$–$Z$ or $R$–$\theta$ plane. Further details of the probe are described in Ref. 6. Plasma light is delivered via fiber optics to a 1 m monochromator (0.05 Å resolution) and imaged with a gated CCD camera (wavelength resolution of 0.074 Å/pixel). The reported $T_i$ values are determined by fitting He II 4686 Å spectra to a single Gaussian convolved with the known instrumental broadening, an example of which is shown in Fig. 2. The approximately 0.2 Å of fine structure in this emission line (slightly less than instrumental broadening) introduces a small systematic error (less than 20%) in the lower range of reported $T_i$ values and can be neglected without altering the physics conclusions of the paper. Time resolution is limited by available light, requiring the width of the CCD gate signal to be $\geq 10 \ \mu$s. Stark broadening is negligible, and IDSP perturbation of the plasma was seen to be minimal after approximately 15 conditioning discharges at the start of each run-day.

IDSP Doppler shifts are also used to calibrate Mach probe measurements of ion flow speed $V_i$ derived from an unmagnetized fluid sheath theory generalized for $T_i \approx T_e$:

$$V_i = \sqrt{\frac{T_e}{T_i}} \sqrt{\frac{T_e}{m_i}} \tanh^{-1} \left( \frac{I_+ - I_-}{I_+ + I_-} \right),$$

(1)

where $I_+$ ($I_-$) is the ion saturation current collected by the upstream (downstream) electrode of the Mach probe. A brief discussion of Mach probes and a derivation of Eq. (1) are given in the Appendix.

### III. IDENTIFICATION OF ION HEATING

#### A. Time evolution of ion temperature

If ions are heated predominantly by the reconnection process, one might expect $T_i$ to increase with time in the reconnection layer when reconnection is driven and $T_i$ to remain constant when reconnection is not driven, provided that all other controllable parameters are unchanged. However, note that failure to observe a rise in $T_i$ does not preclude ion heating because the observed $T_i$ also depends on the rate of ion heat loss from the volume being sampled.
With the IDSP situated in the center of the reconnection region \((R = 35 \rightarrow 40 \text{ cm and } Z = -2.5 \rightarrow 2.5 \text{ cm})\), as shown in Fig. 1(a), \(T_i\) as a function of time was obtained for four cases: with and without reconnection for both null-helicity and co-helicity discharges. “Pull” reconnection is driven when the PF current is allowed to ramp down after reaching its peak, which induces the requisite \(E_\theta\) in the plasma from \(t \approx 250 \rightarrow 280 \mu s\). To suppress reconnection, the PF circuit is shorted out (crowbarred) at the approximate time of peak current so that the current decays over several hundred microseconds and the induced \(E_\theta\) is much smaller. All other controllable parameters, such as capacitor bank voltage and initial gas pressure, were kept constant.\(^{17}\) The \(T_i\) time scan was accomplished by advancing the gate trigger timing for the CCD camera over multiple plasma discharges. The duration over which the CCD gate is open is typically 10 \(\mu s\) (increased to 20 \(\mu s\) early and late in the scans due to limited plasma light).

1. Null-helicity

A substantial rise in \(T_i\) in the center of the reconnection layer, from approximately 6 to 17 eV, is observed when reconnection is driven, and no rise in \(T_i\) is observed when reconnection is not driven, as shown in Fig. 3(top) for null-helicity reconnection. The initial \(T_i \approx 3-5 \text{ eV before } t \approx 245 \mu s\) for both cases is believed to result from the earlier “push” reconnection phase.\(^{18}\) In Fig. 3(top), error bars in the ordinate represent standard deviations in an ensemble of \(T_i\) measurements (5–10 discharges), and error bars in the abscissa represent the CCD gate time. The rise in \(T_i\) is strongly correlated with the magnetic energy dissipation rate, as represented by the value of \(E_\theta j_\theta\) in the center of the reconnection layer, shown in Fig. 3(middle). For the case in which no reconnection is driven, the dissipation rate is smaller by a factor of ten, consistent with \(T_i\) remaining constant. Taking the time integral of \(E_\theta j_\theta\) gives the dissipated magnetic energy per unit volume at the center of the reconnection layer, as shown in Fig. 3(bottom). Note the remarkable agreement in the time evolution of \(T_i\) and the dissipated magnetic energy for both cases, but especially for the case with reconnection.
These observations indicate strongly that reconnection is the energy source for the rise in $T_i$.

The time evolution of $T_e$ is interesting, as shown in Fig. 3(top). The $T_e$ is already at 15 eV early in the discharge, likely due to Ohmic heating associated with plasma formation. The time evolution of $T_e$ starts to diverge around $t \approx 250 \mu s$ for the two cases of with and without reconnection. It might appear surprising at first that $T_e$ decreases when reconnection is driven and stays constant when reconnection is not driven. The most obvious explanation for this effect is if the reconnection layer is pushed toward smaller values of $R$ as reconnection is driven, meaning that the center of the layer with highest $T_e$ sweeps past the stationary triple Langmuir probe. However, examination of $T_e(R,t)$ and $j_\theta(R,t)$ eliminates this explanation, leaving more interesting possibilities. First, the electron confinement characteristics of the two cases may be very different. Second, the most interesting possibility is that electron energy can be converted to ion energy on these short time scales only via non-classical processes which arise when reconnection is driven. An example of such a phenomena due to magnetic fluctuations, which predicts a simultaneous increase in $T_i$ and decrease in $T_e$, has been proposed by Gatto and Terry. However, it must be emphasized that a conclusive explanation of the $T_e$ data awaits further investigation.

A direct ion heating mechanism must be operative if reconnection is in fact responsible for the ion heating. “Direct” means that dissipated field energy is converted to ion energy without requiring classical energy exchange with electrons. In these helium discharges, ions cannot be heated classically by the electrons because the ion-electron energy equipartition time is 400 $\mu s$ or more, and furthermore, the energy gained by electrons during reconnection due to classical Ohmic heating is insufficient by an order of magnitude. One obvious direct ion heating mechanism which would be consistent with the data in Fig. 3 is classical viscous heating by ion flows accelerated due to reconnection. However, classical viscous heating is estimated later in Sec. V to be insufficient in MRX. This leaves only the possibility of non-classical mechanisms.

The direct proportionality between the $T_i$ rise and the dissipated magnetic energy sug-
gests that the magnetic energy dissipation rate is linearly related to the ion heating rate. If true, this would imply that the mechanism allowing reconnection to proceed, i.e. some form of non-classical resistivity, is the same mechanism which is heating the ions. Note that in general the two mechanisms do not have to be the same. This interpretation must be taken cautiously because the rise in $T_i$, which is measured directly, does not have to be proportional to the ion heating rate, which is not measured directly, due to the possible variation of ion loss rates as the dissipation rate changes.

2. Co-helicity

In the co-helicity case, $T_i$ also rises during reconnection, from approximately 3 to 7 eV, as shown in Fig. 4(top). Again, no rise in $T_i$ is observed if reconnection is not driven. The local magnetic energy dissipation rate $E_{\theta j\theta}$ is shown in Fig. 4(middle). In magnitude, it is on average a factor of two smaller than the null-helicity case, consistent with the observed smaller rise in $T_i$. The dissipated magnetic energy as a function of time is shown in Fig. 4(bottom), and the initial rise in $T_i$ is well-correlated with the dissipated magnetic energy. The non-physical dotted section of the curve after $t = 270 \, \mu s$ is a result of the $E_{\theta j\theta}$ term going negative due to the plasma “pinching off” from one flux-core, and this should not be interpreted as part of the main reconnection sequence. The continued rise in $T_i$ after $t = 270 \, \mu s$ is likely due to uncontrolled reconnection and the improved confinement of a spheromak-like configuration which forms.

The time evolution of $T_e$, shown in Fig. 4(top), for the co-helicity case differs from the null-helicity case. First, $T_e \approx 10 \, \text{eV}$ for the co-helicity case compared to $15 \, \text{eV}$ for the null-helicity case. The discussion given previously to account for the difference in $T_e$ after $t \approx 250 \, \mu s$ for null-helicity should apply to the co-helicity case also, except in this case the $T_e$ rise at $t = 270 \, \mu s$ (with reconnection) may be due to the improved confinement of the “pinched-off” spheromak. The difference in $T_e$ before $t \approx 250 \, \mu s$ in this case may be attributed to a very early PF crowbar time of $t = 200 \, \mu s$ for suppressing reconnection.
(compared to \( t = 220 \) µs for null-helicity), meaning there is less early heating due to the push\(^{18}\) reconnection phase for the co-helicity experiments.

3. Comparison of null-helicity and co-helicity

Figure 5 shows the key results from null-helicity and co-helicity together. In the null-helicity case, \( T_i \) rises by an amount \( \Delta T_i \approx 11 \) eV (from \( t = 245 \rightarrow 280 \) µs) and the local dissipated energy at \( t = 280 \) µs is \( 2.1 \) kJ/m\(^3\). In the co-helicity case, which is valid up to \( t = 270 \) µs (due to \( E_\theta \) going negative), \( \Delta T_i \approx 2.5 \) eV (from \( t = 245 \rightarrow 270 \) µs) and the local dissipated energy is \( 0.8 \) kJ/m\(^3\) and \( t = 270 \) µs. These numbers indicate that null-helicity reconnection heats ions more effectively since 2.6 times more dissipated energy resulted in 4.4 times the rise in \( T_i \). Furthermore, the large difference exists despite the fact that co-helicity likely has better ion confinement due to a strong toroidal field.

Stronger ion heating might suggest the increased effect of non-classical dissipation, and this can be investigated with respect to resistivity enhancement over the classical Spitzer value. Figure 6 shows the time evolution of \( E_\theta \) and \( j_\theta \) separately (as well as \( E_\theta j_\theta \)) for null-helicity and co-helicity reconnection.\(^{20}\) A time-averaged effective resistivity, \( \eta^* \equiv E_\theta/j_\theta \), can be determined for the two cases. For the null-helicity case (averaging over \( t \approx 245 \rightarrow 280 \) µs), \( \eta^* \approx (140 \text{ V/m})/(0.4 \text{ MA/m}^2) \approx 3.5 \times 10^{-4} \Omega \cdot \text{m} \) and the classical \( \eta_\perp \approx 2 \times 10^{-5} \Omega \cdot \text{m} \) (\( T_e \approx 15 \) eV, using \( Z_{\text{eff}} = 1 \)) resulting in an enhancement factor of 18. For the co-helicity case (averaging over \( t \approx 250 \rightarrow 270 \) µs), \( \eta^* \approx (50 \text{ V/m})/(0.35 \text{ MA/m}^2) \approx 1.4 \times 10^{-4} \Omega \cdot \text{m} \) and the classical \( \eta_\parallel \approx 1.8 \times 10^{-5} \Omega \cdot \text{m} \) (\( T_e \approx 10 \) eV, using \( Z_{\text{eff}} = 1 \)), resulting in an enhancement factor of 8. The larger resistivity enhancement for null-helicity compared to co-helicity is consistent with the implied stronger heating. The reduced dissipation rate in co-helicity is due to a factor of two reduction in \( E_\theta \), which translates to a slower reconnection speed (since reconnection speed is proportional to \( E_\theta \)). This is consistent with previous findings that co-helicity reconnection is up to a factor of three slower than null-helicity\(^{11,3}\) and counter-helicity reconnection.\(^{21}\)
4. Summary

In these $T_i$ temporal-scan experiments, an increase in $T_i$ during reconnection was identified and causally linked to the presence of reconnection. This is an important result because ion heating due to reconnection had not been identified experimentally before in a rigorous manner. In both null-helicity and co-helicity reconnection, $T_i$ increased when reconnection was driven and stayed constant when reconnection was not driven. The $T_i$ increase correlated well with the dissipation of magnetic energy due to reconnection, showing remarkable proportionality with the dissipated field energy in the null-helicity case. Null-helicity reconnection resulted in a disproportionally higher $T_i$ increase compared to the co-helicity case, indicating stronger ion heating for null-helicity. This is consistent with the fact that null-helicity reconnection had a higher value of resistivity enhancement, which suggests the possible role played by non-classical dissipation mechanisms.

B. Radial profile of ion temperature rise

If ions are heated predominantly by the reconnection process, one would also expect the observable rise in $T_i$ to be largest near the reconnection layer ($R \approx 35 \rightarrow 40$ cm). Indeed, this was seen experimentally.

To determine the spatial profile of the ion heating, the IDSP was scanned in $R$ for two CCD gate times, $t_1 = 250 \rightarrow 260$ µs and $t_2 = 260 \rightarrow 270$ µs; define $\Delta T_i = T_i(t_2) - T_i(t_1)$. Figure 7 shows a peaked profile for $\Delta T_i/T_i(t_1)$ for null-helicity reconnection, suggesting that ion heating occurred in the vicinity of the reconnection layer. The reconnecting field $B_Z$ profile averaged over the same shots is also shown to indicate the location of the reconnection layer (between the “knees” of the $B_Z$ profile, $R \approx 35 \rightarrow 40$ cm). Ion temperature rises approximately 75% in the reconnection layer and not at all elsewhere. Error bars in the ordinate represent one standard deviation in an ensemble of $\Delta T_i$ measurements (approximately 5 discharges at each gate time), and error bars in the abscissa represent the spatial
region from which plasma light is collected by the IDSP.

For the co-helicity case, rise in $T_i$ is small compared to the statistical error between shots, and the observed spatial heating profile is essentially flat (within error bars). Better temporal and spatial resolution than what is currently possible is required to reveal a meaningful spatial profile of $\Delta T_i/T_i$ for co-helicity discharges.

In the above experiment, a spatial correlation between the rise in $T_i$ and the location of the reconnection layer was shown. This is further indication that the reconnection process is directly responsible for the observed rise in $T_i$.

**IV. ION FLOW MEASUREMENTS**

Reconnection is generally expected to accelerate bulk plasma flows. There is evidence for this in solar observational data, magnetospheric satellite data, and in laboratory experiments. For example, in TS-3 at the Univ. of Tokyo,\(^9\) downstream Alfvénic flows were reported and attributed to the strong tension force of reconnected field lines. Alfvénic flows also can develop due to large pressure gradients between the reconnection layer and the far-away downstream region, as considered in the Sweet-Parker model.\(^7,8\) Indeed, reconnection can hardly be mentioned without the expectation of Alfvénic plasma flows. One of the conventionally accepted means of ion heating is via classical viscous damping of these flows. However, measurements of ion flow in MRX indicate that a different scenario is possible, namely that substantial ion heating during reconnection can occur even with sub-Alfvénic flows and small viscous heating.

**A. Downstream flow**

Ion downstream flow speed profiles were measured locally using a Mach probe, the data from which was calibrated using the IDSP.\(^4\) The maximum downstream flow $V_{iz} \approx 8$ km/s, as shown in Fig. 8 for null-helicity reconnection, is equal to $0.2V_A$, where $V_A \approx 39$ km/s ($n_e \approx 5 \times 10^{13}$ cm$^{-3}$, $B \approx 250$ G, and mass of helium). The flow speed is also seen to
increase linearly from 0 to 8 km/s along the layer from $Z = 0 \rightarrow 10$ cm. The pattern of the flow is consistent with 2-D theoretical reconnection models, e.g. Sweet-Parker, but the magnitude of the flow differs substantially (theoretical models generally predict downstream flows at the Alfvén speed $V_A$). Downstream flow for the co-helicity case is similar in profile and magnitude.

The maximum energy density of the measured flow ($\rho V_{iz}^2/2 \approx 11$ J/m$^3$, using $n_e = 5 \times 10^{13}$ cm$^{-3}$) is an order of magnitude smaller than the observed ion thermal energy density increase ($3n\Delta T_i/2 \approx 120$ J/m$^3$, using $n_e = 5 \times 10^{13}$ cm$^{-3}$ which remains roughly constant in time and $\Delta T_i = 10$ eV), implying that the observed ion heating is unlikely to result from thermalization of the outflow. Furthermore, since the flows are stronger at the edges of the layer ($Z = \pm 10$ cm), ion heating due to viscosity could not easily explain the observed ion heating in the center of the reconnection region ($Z = 0$ cm). Estimates of the ion heating due to viscosity are estimated in Sec. V and shown to be small.

The sub-Alfvénic $V_{iz}$ in MRX is consistent with the buildup of high downstream pressure which reduces the $\nabla p$ force along $Z$. Theoretical models generally assume a low ambient downstream pressure, and thus ion flows can be accelerated to $V_A$. In MRX, high downstream electron pressure has been verified by Langmuir probe measurements of $n_e$ and $T_e$ in the downstream region. However, the $Z$-profile of $T_i$ was not measured since the IDSP could not be placed there. The $Z$-profile of $nT_e$ is peaked near $Z = 0$ cm early in the reconnection phase but it becomes flat later in the reconnection phase, mostly due to a rise in $n_e$.

The absence of energetic flows in MRX is an important difference compared to the situations in TS-3$^9$ and SSX,$^{10}$ in which ions are accelerated to $V_A$. The downstream pressures in TS-3 and SSX were likely to be much lower than the pressures in the reconnection layers. In MRX, ion heating occurs without the development of energetic downstream flows, an important new observation. The suppression of Alfvénic downstream flows in MRX due to the buildup of high downstream pressure is an interesting example of how global boundary conditions can affect the local reconnection dynamics.
B. Out-of-plane flow

The out-of-plane reconnection electric field is generally considered to be a likely candidate for particle acceleration to very high energies, especially for highly collisionless events such as solar flares. In MRX with a reconnection electric field $E_\theta \sim 150$ V/m, acceleration via this mechanism is marginal even if particles are allowed to free-stream unimpeded by magnetic fields (it would take more than one classical collision time, 5 $\mu$s for ions and 36 ns for electrons, for either species to reach the 15 eV thermal speed). In reality, the ions are magnetized at the edge of the layer and follow serpentine orbits throughout most of the layer and thus cannot accelerate freely, while the electrons are magnetized down to the $\rho_e$ scale (a thus-far unresolvable 0.4 mm in MRX). Based on these considerations alone, it may be possible to eliminate toroidal ion acceleration as a mechanism for the observed ion heating in MRX. Experimental measurements indeed support this hypothesis. The nature of the toroidal flow, e.g. which species contributes more to the current, may also hold important clues for understanding the equilibrium properties of the current sheet and ultimately of the reconnection process. The Harris equilibrium, which is in good agreement with MRX data with respect to the magnetic field profiles and current sheet thickness, implies a diamagnetic current, but this has not been observed using the present diagnostics.

The toroidal ion flow $V_{i\theta}$ measurements reported here are based on Doppler shifts of the He II 4686 Å ion line emission collected locally by the IDSP. Mach probe measurements of $V_{i\theta}$ indicate relatively good agreement. The IDSP is scanned over several radial positions. At each radial position, 5–10 shots are taken with the sightline at a 45° angle to the $+\theta$ direction, and then this is repeated with the sightline at a 45° angle to the $-\theta$ direction. The averaged difference (of course accounting for the angle of the sightlines) gives the relative Doppler shift, which translates to an absolute toroidal flow velocity. CCD gate time is $t = 250 \rightarrow 270$ μs. It is important to recognize that these measurements have limitations due to averaging effects of the 5 cm light collection volume and the 10 μs CCD gate time, and thus the true local $V_{i\theta}$ may be underestimated.
Shown in Fig. 9 is the radial profile of toroidal flow speed $V_{i\theta}$ for null-helicity reconnection. Vertical error bars represent shot-to-shot variation and horizontal error bars represent the spatial extent of the IDSP. The profile shape is somewhat unexpected, but it has been verified many times by different measurements (Mach probe and chord-averaged spectroscopy). Since there is very little current beyond $R = 40$ cm, these measurements imply that the entire plasma (not just ions) is rotating on the outside. Most importantly, note that the magnitude of $V_{i\theta}$ in the layer ($R \approx 35 \rightarrow 40$ cm) is only one-tenth the ion thermal speed ($\approx 20$ km/s for 15 eV helium ions) and does not account for the observed ion heating via classical viscous thermalization. The co-helicity case is quite different, as shown in Fig. 10. The flow profile is symmetric and peaked slightly outside the current sheet. The peak value of $V_{i\theta}$ in the layer is higher than for the null-helicity case but is still only a few km/s at most ($< 1$ eV).

The important conclusion of the toroidal flow measurements is that the observed ion heating does not originate from energetic toroidal ion flows in the layer. In the reconnection layer, the instantaneous energy density of these flows ($\rho V_{i\theta}^2 / 2 \lesssim 1.5 \text{ J/m}^3$) is two orders of magnitude smaller than the observed ion thermal energy density ($\approx 120 \text{ J/m}^3$, as given in the previous section) and, as in the downstream flow case, classical viscous heating by these flows is insufficient to account for the ion heating. Regarding the toroidal current, the magnitude of the measured $V_{i\theta}$ implies that the current is not diamagnetic, which would predict an ion drift speed of $j_{\theta} / 2ne \approx 30$ km/s (using $j_{\theta} = 0.5 \text{ MA/m}^2$ and $n_e = 5 \times 10^{13} \text{ cm}^{-3}$). Below, speculation regarding the apparent contradiction is given.

The value of peak toroidal current density in the layer, $j_{\theta} \approx 0.5 \text{ MA/m}^2$, derived from $B_z$ measurements implies a relative drift between ions and electrons of $V_D \equiv V_{i\theta} - V_{e\theta} = j_{\theta} / ne \approx 62$ km/s (using $n_e = 5 \times 10^{13} \text{ cm}^{-3}$). The toroidal flow measurements indicate an ion drift speed much less than the diamagnetic speed. This deviation, as well as the difference in $V_{i\theta}$ between null-helicity and co-helicity reconnection, might be explained by the different radial electric fields $E_R$ present in the two cases. Langmuir probe floating potential $V_f$ measurements are shown in Figs. 11 and 12 for null-helicity and co-helicity, respectively. In
the null-helicity case, \( V_f \) decreases monotonically with \( R \), yielding \( E_R \sim 1000 \text{ V/m} \). The resulting toroidal \( E_R \times B_Z \) drift \([\sim (1000 \text{ V/m})/(0.03 \text{ T}) = 33 \text{ km/s}] \) is the same order of magnitude as the strong ion rotation outside the layer, as seen in the data of Fig. 9, and can cancel the contribution of ion diamagnetic drift in the layer. This possibly explains why the diamagnetic drift speed is not observed in the \( V_{i\theta} \) measurements. In contrast, the \( V_f \) profile for co-helicity reconnection is essentially flat, meaning there is no macroscopic radial electric field. In this case, \( V_{i\theta}(R) \) is fairly symmetric and limited in magnitude. In the co-helicity case, the current would not be expected to be diamagnetic since there is a strong flow component parallel to the toroidal field. It is unclear what ion drift speed ought to be expected in the co-helicity case. Ongoing investigations using Langmuir probe arrays will hopefully lead to better understanding of the equilibrium properties of the MRX current sheet. It is particularly curious that \( V_{i\theta} \) and \( V_f \) profiles differ from the Harris model while the magnetic field profile and current sheet thickness do not. Note, however, that the Harris theory is also valid in the ion rest frame, in which case the ion flows are zero but the magnetic field profiles remain unchanged.

\section*{V. ION ENERGY BALANCE DURING RECONNECTION}

In this section, an energy balance for ions during null-helicity reconnection is considered in detail. The goal is to determine a lower bound on the energy gained by ions during the reconnection process and how much of that energy is converted non-classically. The energy balance equation used here is based on Eq. (1.23) of Braginskii:  

\[
\frac{3}{2} \frac{\partial (nT_i)}{\partial t} + nT_i \nabla \cdot \mathbf{V} + \nabla \cdot \left( \frac{3}{2} nT_i \mathbf{V} \right) + \nabla \cdot \mathbf{q} = Q_{i-n} + Q_{\text{vis}} + Q_{i-e} + Q_{\text{noncl}},
\]

where the terms on the left hand side (LHS) are (from left to right) rise in ion thermal energy density, a compression term including work done by pressure, convective heat loss, and conductive heat loss. The terms on the right hand side (RHS) are sinks and sources of energy for the ions, including (from left to right) energy lost to neutrals, viscous heating,
heating due to electron-ion collisions, and any non-classical ion heating mechanisms (e.g., due to wave-particle interactions). The second and third terms on the RHS represent energy due to classical heating. The total ion heating is some fraction of the reconnected field energy, which is known from $E_\theta j_\theta$ derived from magnetic probe measurements. Each term of Eq. (2) will be estimated based on experimental data wherever possible. Note that $Q_{i-e}$ can be neglected since the characteristic classical ion-electron energy equilibration time is more than 400 $\mu$s and thus irrelevant on the time scale of the reconnection process (30 $\mu$s).

The results show that a substantial fraction (more than half) of the reconnected field energy is converted to ion energy, and mostly due to non-classical mechanisms. These results differ from reported TS-3 results\(^9\) in which ion heating was attributed predominantly to viscous damping of Alfvénic ion flows, and from classical MHD reconnection models in which ions are heated both by viscosity and by energy exchange with Ohmically heated electrons.

In calculating an energy balance, a finite volume and time duration must be defined. Here, the energy balance is considered for a given volume $V = 5.9 \times 10^{-3}$ m$^3$, which is the area monitored by the IDSP ($R = 35 \rightarrow 40$ cm and $Z = -2.5 \rightarrow 2.5$ cm) revolved around the axis of symmetry. The volume $V$ is essentially a toroid coinciding with the center of the toroidal current sheet, with the area monitored by the IDSP as its toroidal cross section. The time duration $\Delta t$ is defined as $t_1 = 245 \rightarrow t_2 = 265$ $\mu$s, which is during the pull reconnection phase. The data are from a set of null-helicity discharges with the same parameters\(^{17}\) as the ones in Figs. 3 and 7. Note that all the calculations in this section invoke axisymmetry since measurements are known only at one toroidal position.

### A. Reconnected field energy

Before examining the terms in Eq. (2), the energy released due to reconnection will be calculated first. The dissipated magnetic field energy can be calculated according to

$$W_{\text{rec}} = \int_{t_1}^{t_2} \iint_V \mathbf{E} \cdot \mathbf{j} \, d^3V \, dt \approx \int_{t_1}^{t_2} \iint_V E_\theta j_\theta \, d^3V \, dt,$$

(3)
which represents the total energy due to reconnection available to heat ions (and electrons). The toroidal electric field $E_\theta$ and current density $j_\theta$ as a function of space and time are shown in Figs. 13 and 14, respectively.

Note that $E_\theta$ is relatively uniform in space, while $j_\theta$ is peaked near $R = 37$ cm. Both quantities increase in time as pull reconnection proceeds and then reach a short flat-top around $t = 260 \mu s$ before starting to decrease. Energy dissipation is clearly strongest in the reconnection layer where $j_\theta$ is concentrated. Using the $E_\theta$ and $j_\theta$ data in Eq. (3) gives $W_{\text{rec}} \approx 4.8 \pm 0.7$ J. It should be noted that the volume $V$, is only a fraction of the total reconnection volume, and that the total energy dissipated in the entire reconnection volume during $\Delta t$ is estimated to be on the order of 30 J.

**B. Increase in ion thermal energy**

The increase in ion thermal energy, the first term on the LHS of Eq. (2), in volume $V$ during reconnection is calculated as follows,

$$\Delta W_{\text{th,i}} \equiv \frac{3}{2} [\bar{n}(t_2)T_i(t_2) - \bar{n}(t_1)T_i(t_1)]V \approx 0.5 \pm 0.2 \text{ J},$$

where $\bar{n}(t_1) \approx \bar{n}(t_2) \approx 5 \pm 1 \times 10^{13}$ cm$^{-3}$ is the density averaged over area $A$, and $T(t_1) \approx 6 \pm 1$ eV and $T(t_2) \approx 13 \pm 2$ eV (see Fig. 3). Note that $\Delta W_{\text{th,i}}$ is only the remnant ion thermal energy in $V$ and does not include ion heat loss during $\Delta t$. The rise in ion thermal energy is due predominantly to a rise in $T_i$, as the density in the layer remains fairly constant during $\Delta t$, as shown in Fig. 15.

As an aside, an additional set of experiments in which the firing voltage was varied is reported here. A range of $\Delta W_{\text{th,i}}$ and $W_{\text{rec}}$ were obtained in this experiment. Figure 16 shows that $\Delta W_{\text{th,i}}$ scales with $W_{\text{rec}}$, i.e. $T_i$ rises more as more energy is released due to reconnection. Each data point represents the average of 5 plasma shots at a given firing voltage, and error bars (from data point to extreme) represent one standard deviation in the shot-to-shot scatter. This is consistent with the conclusion from Sec. III that the observed rise in $T_i$ is causally linked to reconnection.

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C. Compression

The second term on the LHS of Eq. (2) is a compression term including the work done by pressure forces. It can be estimated approximately as \( \langle nT_i(\partial V_R/\partial R + \partial V_Z/\partial Z) \rangle V \Delta t \) (the \( V_R/R \) term of \( \nabla \cdot V \) averages to nearly zero over \( V \) and is neglected), where the brackets indicate a spatial average over \( V \) and temporal average over \( \Delta t \). Using values of \( n \approx 5 \pm 1 \times 10^{13} \text{ cm}^{-3}, T_i \approx 10 \pm 1 \text{ eV}, \partial V_R/\partial R \approx -(4.8 \text{ km/s})/(0.05 \text{ m}) = -9.6 \pm 1 \times 10^4 \text{ s}^{-1}, \) and \( \partial V_Z/\partial Z \approx (8 \text{ km/s})/(0.1 \text{ m}) = 8 \pm 2 \times 10^4 \text{ s}^{-1}, \) the compression term is \( W_{\text{compression}} \approx -0.16 \pm 0.31 \text{ J}. \) The negative value means that there is some ion heating in volume \( V \) due to compression. The large relative error arises due to the subtraction of the velocity gradient terms, which results in a small number with large relative error. It will be shown that this compression term is small compared to the ion loss terms, and thus the rough estimate adopted above is justified.

D. Ion energy losses

Since volume \( V \) is an open system, a complete ion energy balance must consider ion energy losses due to convective and conductive heat loss. The energy lost must be accounted for since their original source was the dissipated magnetic energy. The heat conduction estimates in this sub-section are based on classical transport. Therefore, they likely represent a lower bound on ion energy loss since the transport is almost certainly not classical. It is important to emphasize that this would only lead to an underestimate of energy conversion to ions. In general, ions may also lose energy to neutrals, and this is discussed.

1. Convection

Convective heat loss, the third term on the LHS of Eq. (2), out of \( V \) during \( \Delta t \) can be estimated as the heat convected out of \( V \),

\[
W_{\text{out}} \approx \left[ \frac{3}{2} n(\text{layer})T_i(\text{layer}) \right] \times (\text{volume of plasma out}),
\]  

(5)
minus the heat convected into $V$,

$$W_{\text{in}} \approx \left[ \frac{3}{2} n(\text{upstream})T_i(\text{upstream}) \right] \times (\text{volume of plasma in}),$$

(6) during $\Delta t$. The volumes of plasma flowing out of and into $V$ are determined based on the outflow speed $V_{iZ} \approx 3 \pm 0.8$ km/s and inflow speed $V_{iR} \approx 3 \pm 0.4$ km/s, respectively. At the boundaries of volume $V$, the two speeds are approximately equal and thus the volumes convected in and out are also equal ($1.4 \times 10^{-2}$ m$^3$). The density in the layer $n(\text{layer})$ is approximately $5 \pm 1 \times 10^{13}$ cm$^{-3}$, and the average upstream density is approximately $3 \pm 0.4 \times 10^{13}$ cm$^{-3}$. The layer $T_i$ is approximately $12 \pm 0.9$ eV and the average upstream $T_i \approx 10 \pm 1.3$ eV. These numbers combine to yield a total ion energy loss due to convection out of volume $V$ during time $\Delta t$ of $W_{\text{convection}} = W_{\text{out}} - W_{\text{in}} \approx 1.0 \pm 0.7$ J.

2. Conduction

Ion heat loss due to thermal conduction is estimated using classical transport. This estimate likely represents a lower bound because (1) $T_i$ gradients may be underestimated due to the time resolution of the measurements, and (2) enhanced non-classical transport mechanisms are unknown but are likely to exist. Ion energy loss due to classical conduction can be written as

$$\nabla \cdot q = -\nabla_\perp \cdot (\kappa_\perp \nabla_\perp T_i) - \nabla_\parallel \cdot (\kappa_\parallel \nabla_\parallel T_i) + \nabla \cdot \left[ \frac{5cnT_i}{2eB}(B/B \times \nabla T_i) \right],$$

(7) where the last term on the RHS vanishes since $(B/B \times \nabla T_i)$ has only a $\theta$ component and $\partial/\partial \theta = 0$, and $\kappa_\perp$ and $\kappa_\parallel$ are the perpendicular and parallel ion thermal conductivities, respectively,

$$\kappa_\perp = \frac{2nT_i}{m_i \omega_0^2 \tau_i} \approx 2.8 \pm 0.6 \times 10^{22} \frac{1}{\text{m} \cdot \text{s}}$$

(8)

$$\kappa_\parallel = \frac{3.9nT_i \tau_i}{m_i} \approx 11.0 \pm 2.2 \times 10^{22} \frac{1}{\text{m} \cdot \text{s}}$$

(9)

($T_i$ in eV and all other variables in cgs units). Average values during $\Delta t$ have been used: $n = 5 \pm 1 \times 10^{13}$ cm$^{-3}$, $T_i = 10 \pm 1$ eV, and $B = 250$ G. The total surface area of $V$ is
0.48 m², and it is equally divided between being intercepted by $B_Z$ in the perpendicular and parallel directions. Using these values and $\nabla_{\perp} T_i \approx (2 \pm 1 \text{ eV})/(5 \text{ cm})$ and $\nabla_{\parallel} T_i \approx (1 \pm 0.5 \text{ eV})/(10 \text{ cm})$, the ion heat loss due to classical transport is estimated to be $W_{\text{conduction}} \approx 1.7 \pm 0.7 \text{ J}$.

3. Collisions with neutrals

Ion-neutral collisions can be another energy loss channel for heated ions. In MRX regimes, charge exchange is expected to be the dominant ion-neutral interaction. By any reasonable estimate, the plasma inside the current sheet with $T_e > 15 \text{ eV}$ should be better than 99% ionized. The cross section for He–He$^+$ charge exchange for 10 eV ions is $\langle \sigma v \rangle \approx 4 \times 10^{-9} \text{ cm}^3/\text{s}$. The ion density is approximately $5 \times 10^{13} \text{ cm}^{-3}$. A concentration of neutrals equal to 1% of the measured plasma density would result in a charge exchange time of 500 $\mu$s, completely negligible on the reconnection time scale of 30 $\mu$s. It would take more than a 50% concentration of neutrals in the reconnection layer to bring down the charge exchange time to a more relevant 10 $\mu$s. A 50% neutral concentration in the presence of 15 eV electrons is highly unlikely. Therefore, ion energy loss to neutrals is neglected. Note that ion-neutral energy loss would only increase the estimate of ion heating due to reconnection.

E. Classical viscous heating

Classical heating per unit volume due to viscosity in the absence of a magnetic field (justified since ions are unmagnetized in the layer) is

$$Q_{\text{vis}} = \eta_0 W_{\alpha\beta} \frac{\partial V_\alpha}{\partial x_\beta},$$

(10)

where $\eta_0 = 0.96 n T_i \tau_i \approx 1.8 \times 10^{-4} \text{ J} \cdot \text{s/m}^3$ (using $n = 5 \times 10^{13} \text{ cm}^{-3}$ and $T_i = 10 \text{ eV}$), and the rate-of-strain tensor is given by
\[ W_{\alpha\beta} = \frac{\partial V_\alpha}{\partial x_\beta} + \frac{\partial V_\beta}{\partial x_\alpha} - \frac{2}{3} \delta_{\alpha\beta} \nabla \cdot V. \] (11)

The non-zero velocity gradients averaged over volume \( V \) are

\[
\begin{align*}
\frac{\partial V_R}{\partial R} &\approx -\frac{4.8 \text{ km/s}}{0.05 \text{ m}} = -9.6 \times 10^4 \text{ s}^{-1} \quad (12) \\
\frac{\partial V_{IZ}}{\partial Z} &\approx \frac{8 \text{ km/s}}{0.1 \text{ m}} = 8.0 \times 10^4 \text{ s}^{-1} \quad (13) \\
\frac{\partial V_{IZ}}{\partial R} &\lesssim \frac{3 \text{ km/s}}{0.05 \text{ m}} = 6.0 \times 10^4 \text{ s}^{-1} \quad (14) \\
\frac{\partial V_{i\theta}}{\partial R} &\lesssim \frac{2 \text{ km/s}}{0.05 \text{ m}} = 4.0 \times 10^4 \text{ s}^{-1}, \quad (15)
\end{align*}
\]

where \( V_{IZ} \) and \( V_{i\theta} \) are based on Mach probe and IDSP measurements, and \( V_{iR} \) is based on taking the \( E_\theta B_Z \) inflow velocity outside the layer and dividing by the scale length of the layer. Therefore the classical viscous heating per unit volume is

\[
Q_{\text{vis}} \approx \eta_0 \left( W_{RR} \frac{\partial V_R}{\partial R} + W_{ZZ} \frac{\partial V_{IZ}}{\partial Z} + W_{ZR} \frac{\partial V_{IZ}}{\partial R} + W_{\theta R} \frac{\partial V_{i\theta}}{\partial R} \right) \approx 6.5 \pm 2.2 \times 10^6 \text{ J/s} \cdot \text{m}^3 \quad (16)
\]

Using these numbers, the maximum ion energy gained via viscous damping of ion flows is

\[
W_{\text{vis}} \approx Q_{\text{vis}} V \Delta t = 0.8 \pm 0.3 \text{ J}.
\]

F. Non-classical ion heating

The important terms of Eq. (2) are summarized in Table I, which gives a quantitative description of energy conversion to ions based on experimental data. The numbers show that 65 ± 21% of the dissipated field energy is converted to ion energy. Now the question is how much of that energy was converted via non-classical mechanisms. As mentioned before, classical heating due to ion-electron collisions is neglected due to the long ion-electron energy partition time. In any case, energy available due to Ohmic heating is insufficient, estimated to be \( \eta_{Sp} j_\theta^2 V \Delta t \approx 0.2 \text{ J} \) (using \( \eta_{Sp} = 2 \times 10^{-5} \text{ \Omega m} \) and \( j_\theta \approx 0.3 \text{ MA/m}^2 \)). Note that this is only 4% of the total dissipated magnetic energy. The other classical mechanism is viscous heating by the ion flow, which was estimated in the previous section to be 0.8 ± 0.3 J. Subtracting \( W_{\text{viscosity}} \) from \( W_{\text{ions}} \) leaves 2.3 ± 1.0 J of energy which must have been converted
to ion energy via non-classical mechanisms. Thus, $48 \pm 21\%$ of the dissipated magnetic energy was converted to ions non-classically. The various energy components are illustrated in Fig. 17.

In this section, the energy budget of ions was considered carefully. The total energy released due to reconnection was calculated from experimental data and compared to the different components of ion energy, including observable rise in ion thermal energy and ion heat loss due to convection and conduction. The energy budget showed that some 48% of the dissipated magnetic energy was converted to ions non-classically. This finding on non-classical ion heating during reconnection, based on experimental measurements, has significant implications for reconnection research since the basic mechanisms which are responsible for enhanced reconnection rates and the energy conversion process are still poorly understood.

VI. DISCUSSION

A. Correlation of ion heating with “resistivity enhancement”

In the previous section, it was established that ions must have been heated via non-classical dissipation mechanisms. In this section, experimental data is given suggesting a correlation between ion heating and resistivity enhancement. Resistivity enhancement is defined as the ratio of the measured plasma resistivity $\eta^*$ to the classical Spitzer resistivity $\eta_{Sp}$. The relative importance of non-classical versus classical dissipation is embodied in the enhancement factor $\eta^*/\eta_{Sp}$, which has been shown to increase as collisionality decreases.$^{12}$ One possible explanation for this effect is that as the plasma becomes more collisionless, wave fields can scatter current-carrying particles, increasing $\eta^*$ (and possibly also heating ions more efficiently). Generally, enhanced values of resistivity might be expected to be an electron effect since they are expected to carry most of the current. However, in null-helicity reconnection, the current is perpendicular to the field, and it is not clear that electrons
should be carrying most of the current. In fact, the current may be largely diamagnetic.\textsuperscript{14}

By scanning the discharge voltage\textsuperscript{26} in a set of null-helicity experiments, $\Delta W_{\text{th,i}}/W_{\text{rec}}$ and $\eta^*/\eta_{\text{Sp}}$ are varied. Increasing the discharge voltage increases reconnection and thus increases ion heating. The scaling of the quantities with each other can provide insight into the nature of the non-classical ion heating mechanisms. Shown in Fig. 18 is $\Delta W_{\text{th,i}}/W_{\text{rec}}$ versus $\eta^*/\eta_{\text{Sp}}$. The fraction $\Delta W_{\text{th,i}}/W_{\text{rec}}$ increases from approximately 4\% to 14\% as $\eta^*/\eta_{\text{Sp}}$ increases from 8 to 15, as shown in Fig. 18. (Note again that $\Delta W_{\text{th,i}}$ is not the total energy gained by ions but only the remnant ion thermal energy in volume $V$ not including ion heat loss.) Although the error bars are sizable, the trend between the energy converted to ions and the resistivity enhancement is clear. The subtle, and perhaps unexpected, hint provided by Fig. 18 is that the (non-classical) mechanism determining the enhanced reconnection rate (and hence increased reconnected field energy) may also be responsible for channeling the reconnected field energy to the ions. The details, however, would also depend on the ion heat loss characteristics.

\section*{B. Possible ion heating mechanisms}

The experimental data have indicated the existence of non-classical ion heating, non-classical resistivity, and possibly a correlation between the two. The obvious next step experimentally, presently underway, is to investigate the frequency spectrum of turbulent fluctuations with the possibility of identifying relevant modes which can explain the enhanced resistivity and possibly also the ion heating. Theoretically, much effort has been devoted to addressing these issues and a full review is beyond the scope of this paper.\textsuperscript{27}

There are several free energy sources in the MRX current sheet which can drive unstable fluctuations, including a large density gradient and cross field current. Turbulence driven by unstable modes could provide a mechanism to produce resistivity and ion heating in current sheets. One instability which has been extensively investigated theoretically in the context of current sheets is the lower-hybrid drift instability.\textsuperscript{28,29} It is a high frequency
(\Omega_i \ll \omega \ll \Omega_e) mode driven unstable by cross-field current and associated density gradients, and it persists in the regime \( T_e \lesssim T_i \) as required in MRX. The mode is strongly growing in \( \rho_i \)-scale density gradients and has a wavelength near the electron gyroradius and a phase velocity near the ion thermal speed. The value of the phase velocity allows for the possibility of resonant interaction with and heating of the ion population in the MRX current sheet. It is well known that the mode is linearly stabilized at high plasma \( \beta \) and should exist with appreciable amplitude only at the edges of the current sheet, while non-classical resistivity and ion heating are seen in the center of the MRX current sheet. However, strong radial transport produced by this mode could result in enhanced resistivity. Measurements of fluctuations in MRX show the presence of strong fluctuations in the lower-hybrid frequency range at the edges of the current sheet, and dispersion of these fluctuations is consistent with the lower-hybrid drift mode. Studies of the relationships of these modes to the reconnection process are currently underway.

Another possible ion heating mechanism manifests itself in reconnection simulations based on a 2.5-D hybrid code (kinetic ions, fluid electrons) which includes Hall dynamics and electron inertia. In this simulation, a two-scale structure develops in the reconnection layer owing to the Hall effect, which allows ions and electrons to decouple on scale lengths shorter than \( c/\omega_{pi} \), the scale at which ions become unmagnetized. (MRX data also shows a current sheet thickness of \( c/\omega_{pi} \).) The electrons continue inward into the layer until the \( c/\omega_{pe} \) scale at which they finally become unmagnetized and the flux-freezing constraint is finally broken. A self-consistent electric field arises between the \( c/\omega_{pi} \) and \( c/\omega_{pe} \) scales due to the charge separation and can accelerate inflowing ions up to the Alfvén speed. The counter-streaming ions which come in from both sides of the layer mix and appear to be heated instantaneously. This would seem to be consistent with the immediate rise in \( T_i \) observed in the center of the reconnection layer in MRX. However, when 3-D effects are included in the simulations, the thin electron layer breaks up turbulently, and it is not clear if this ion heating mechanism would survive in a physical reconnection layer. Furthermore, the simulations are collisionless and may not model accurately the dominant dissipation processes in the MRX reconnection.
layer. An electrostatic probe array is being constructed in order to investigate the electric field structure across the reconnection layer and the likelihood of two-fluid effects playing a role in MRX.

VII. SUMMARY

In this work, local ion temperature was measured in a well-characterized reconnection layer. A rise in $T_i$ during reconnection was identified, for both null-helicity ($6 \rightarrow 17$ eV) and co-helicity ($3 \rightarrow 7$ eV) reconnection. Furthermore, the $T_i$ rise was causally linked to the reconnection process, i.e. the $T_i$ rise occurred only when reconnection was driven, and the magnitude of the rise was in direct proportion to the amount of magnetic field energy dissipated. Additionally, the rise in $T_i$ was shown to be localized spatially in the region of the reconnection current sheet and magnetic field reversal. These results collectively could be interpreted as the first clear experimental demonstration of ion heating due to magnetic reconnection.

Local downstream ion flow measurements showed the flows to be small, at most 25% of $V_A$ ($0.25V_A \approx 8$ km/s). This observation is consistent with the high downstream pressure observed in MRX, postulated to arise due to fast pressure equalization in the finite downstream volume. The small flow speeds of MRX are in direct contrast to the results of TS-3$^9$ and SSX$^{10}$ as well as to classical MHD reconnection theories which ignore the effects of downstream pressure. Toroidal flows were also measured and shown to be small (a few km/s at most) and not a possible energy source for the observed ion heating. The toroidal flow measurements are not consistent with the Harris sheet diamagnetic ion drift, but the discrepancy may be related to the existence of a radial electric field which causes an $E \times B$ drift modifying the diamagnetic drift. Further investigations are needed to resolve the discrepancy.

An energy balance for the reconnected field and ion energy was considered. The analysis showed that substantial ion heating occurred and that approximately 65% of the reconnected
field energy was converted to ion thermal energy. At most, 17% of the dissipated field energy would have been converted to ion energy via classical viscosity, meaning that 48% of the energy was converted non-classically. Finally, there is a hint of stronger ion heating with increased resistivity enhancement, suggesting a relationship between the non-classical mechanisms responsible for enhanced resistivity and ion heating.

The exact mechanism(s) for the heating and associated “anomalous” resistivity is still an open question and the subject of ongoing research, both experimentally and theoretically. Conversely, the identification of non-classical ion heating on MRX should impact current thinking on the possible non-classical micro-physics in the reconnection layer.

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APPENDIX A: MACH PROBE DATA ANALYSIS

A Mach probe collects ion saturation current $I_{\text{sat}}$ on each of two oppositely facing electrodes. Intuitively, it is reasonable to expect that the difference between the collected currents should reveal information about the average ion drift speed $V_i$ past the probe. The experimental quantity measured is

$$K \equiv \frac{I_+ - I_-}{I_+ + I_-},$$

where $I_+$ and $I_-$ are the (ion saturation) currents collected by the upstream and downstream electrodes, respectively. By considering the appropriate probe sheath physics, a theoretical relationship between $K$ and $V_i$ can be postulated.
A large number of Mach probe papers exist in the literature, and a review of them will not be attempted here. The important parameters to consider are $a/\lambda_D$, $T_i/T_e$, and $\rho_i/a$, where $a$ is the characteristic probe electrode dimension and $\lambda_D$ is the Debye length. MRX plasma and Mach probe parameters fall into the regime $a/\lambda_D \gg 1$, $T_i \gtrsim T_e$, and $\rho_i/a \gg 1$, in which there is no general, rigorous theory to predict $I_{sat}$. However, the model of Hudis and Lidsky can be used as a starting point. This model is based on the Bohm sheath model, in which $T_i \ll T_e$, but modified to include a net ion drift $V_i (\equiv \sqrt{(T_e + T_i)/m_i})$ at the sheath edge. The ion saturation current in this case can be approximated by

$$I_{\pm} = \exp(-1/2)n_{i0}eAC_s \exp\left(\frac{m_i V_i^2}{2T_e}\right) \exp\left(\pm \sqrt{\frac{m_i I_i V_i}{T_e}}\right),$$  \hspace{1cm} (A2)$$

where $n_{i0}$ is the density far from the probe and $A$ is the electrode surface area. Numerical solutions for $I_{sat}$ have shown that the Bohm sheath model is not strongly dependent on $T_i$ for $T_i \lesssim T_e$. Since $T_i \sim T_e$ in the experiments for which ion flow speeds are measured, Eq. (A2) can still be used. Substituting $I_{\pm}$ into Eq. (A1) yields

$$V_i = \sqrt{\frac{T_e}{T_i}} \sqrt{\frac{T_e}{m_i}} \tanh^{-1} K,$$  \hspace{1cm} (A3)$$

the same expression given in Eq. (1).

It must be emphasized that, as a general rule, theoretical expressions for electrostatic probes such as Eq. (A3) are usually not accurate to better than a factor of two. This is not surprising given the non-rigorous justifications relied upon in the above discussion. The trends (spatial profile and temporal evolution) which are measured, however, are still meaningful. Ideally, the absolute value of Mach probe results should be calibrated with an independent measurement. It is found that Mach probe measurements based on Eq. (A3) agree with Doppler shifts of ions obtained by spectroscopy to better than 50%. Thus, $V_i$ values reported in this paper are based on Eq. (A3) but include an empirical calibration factor of 0.75.
REFERENCES


TF/PF bank voltages=13/11 kV; TF/PF bank capacitances=120 µF; TF/PF trigger=180/270 µs; TF/PF crowbar (w/reconnection)=270/290 µs; TF/PF crowbar (w/o reconnection): 270/220 (null-helicity), 270/200 (co-helicity); initial He gas pressure=6 mT.

During plasma formation, the current is increasing and private flux is reconnected into public flux, establishing a reconnection layer elongated in the $R$ direction. This is push reconnection.

The negative values of $E_\theta$ and $j_\theta$ before $t \approx 248 \mu s$ for co-helicity are due to the end of the push reconnection phase in which the induced $E_\theta$ is in the other direction. The negative $E_\theta$ after $t = 270 \mu s$ is likely due to the plasma pinching off from one flux-core.

The high downstream pressure in MRX may arise due to fast pressure equalization along field lines which connect the reconnection region to the finite-volume downstream regions near the flux-cores which probably have better confinement. The thermal transit time of 15 eV electrons along this path is on the order of 1 µs, meaning that the buildup of downstream pressure due to recirculating electrons over the 20–30 µs duration of pull reconnection is plausible.

TF/PF bank voltages=11/9, 12/10, 12.5/10.5, 13/11, 14/12 kV.
27 see e.g. F. V. Coroniti, J. Geophys. Res. 90, 7427 (1985).


TABLES

TABLE I. Ion energy budget between $t = 245 \rightarrow 265 \mu s$ in volume $V$, showing that some 65\% of the dissipated magnetic energy $W_{rec}$ is converted to ions $W_{ions}$.

<table>
<thead>
<tr>
<th>process</th>
<th>energy (J)</th>
</tr>
</thead>
<tbody>
<tr>
<td>$W_{rec}$</td>
<td>4.8 ± 0.7</td>
</tr>
<tr>
<td>total $W_{ions}$</td>
<td>3.1 ± 1.0</td>
</tr>
<tr>
<td>$\Delta W_{th,i}$</td>
<td>0.5 ± 0.2</td>
</tr>
<tr>
<td>$W_{compression}$</td>
<td>-0.14 ± 0.28</td>
</tr>
<tr>
<td>$W_{convection}$</td>
<td>1.0 ± 0.7</td>
</tr>
<tr>
<td>$W_{conduction}$</td>
<td>&gt; 1.7 ± 0.7</td>
</tr>
<tr>
<td>$W_{vis}$</td>
<td>0.8 ± 0.3</td>
</tr>
</tbody>
</table>
CAPTIONS

FIG. 1. (a) A schematic of the upper-half plane of MRX, showing the flux cores, position of the IDSP, and the coordinate system. (b) A schematic illustrating motion of plasma and contours of constant poloidal flux and formation of the pull reconnection layer.

FIG. 2 An example of the He II 4686 Å spectral line, as measured by the IDSP, fitted with a convolution of a Gaussian and the known instrumental broadening.

FIG. 3 Time evolution of (top) $T_i$ and $T_e$, (middle) local heating rate $E_{\theta j\theta}$, and (bottom) dissipated magnetic energy $\int_{240\mu s}^{t}(E_{\theta j\theta})dt'$ per unit volume, all in the center of the reconnection layer for null-helicity discharges (IDSP located at $R = 37.5$ cm and $Z = 0$ cm).

FIG. 4 Time evolution of (top) $T_i$ and $T_e$, (middle) local heating rate $E_{\theta j\theta}$, and (bottom) energy dissipated $\int_{240\mu s}^{t}(E_{\theta j\theta})dt'$ per unit volume, all in the center of the reconnection layer for co-helicity discharges (IDSP located at $R = 37.5$ cm and $Z = 0$ cm). The non-physical decrease in dissipated energy (represented by the dotted line in the bottom panel) is a result of $E_{\theta}$ going negative due to the plasma “pinching off” from one flux-core.

FIG. 5 Time evolution of (top) $T_i$ with the IDSP placed in the center of the reconnection layer ($R = 37.5$ cm and $Z = 0$ cm) and (bottom) dissipated magnetic energy for both null-helicity and co-helicity reconnection.

FIG. 6 Time evolution of (top) $E_{\theta}$, (middle) $j_{\theta}$, and (bottom) $E_{\theta j\theta}$ for null-helicity and co-helicity, all in the center of the reconnection layer.

FIG. 7 Radial profiles of (solid line) the relative rise in $T_i$ from $t = 255 \to 265$ $\mu$s ($R = 37.5$ cm, $Z = 0$ cm) and (dashed line) the reconnecting field $B_Z$ during the same time (also at $Z = 0$ cm) for null-helicity discharges (shots 11614–11659).
FIG. 8 Mach probe measurements of $V_{iZ}$ as a function of $R$ at $Z = 10$ cm and as a function of $Z$ at $R = 38$ cm averaged over $t = 250 \rightarrow 270$ µs.

FIG. 9 IDSP measurement of toroidal ion flow $V_{i\theta}$ at $Z = 0$ cm and $t = 250 \rightarrow 270$ µs for null-helicity reconnection (shots 10981–11097).

FIG. 10 IDSP measurement of toroidal ion flow $V_{i\theta}$ at $Z = 0$ cm and $t = 250 \rightarrow 270$ µs for co-helicity reconnection (shots 11343–11428), showing a more symmetric profile.

FIG. 11 Radial profile of Langmuir probe floating potential (with respect to machine ground) at three times during null-helicity reconnection (shots 11010–11133).

FIG. 12 Radial profile of Langmuir probe floating potential (with respect to machine ground) at three times during co-helicity reconnection (shots 11386–11424).

FIG. 13 Surface plot of null-helicity $E_{\theta}$ as a function of radius and time, averaged over $Z = -2.5 \rightarrow 2.5$ cm (from 90-channel probe measurements, averaged over shots 12233–12266).

FIG. 14 Surface plot of null-helicity $j_{\theta}$, which is peaked in the reconnection layer, as a function of radius and time at $Z = 0$ cm (from high-resolution 1-D probe measurements, averaged over shots 12233–12266).

FIG. 15 Observed increase in ion thermal energy $\Delta W_{th,i}$ versus reconnected field energy $W_{rec}$ for null-helicity discharges with varying firing voltages.
FIG. 16 Time evolution of $n_e$ measured by triple Langmuir probe at $R = 37.5$ cm and $Z = 0$ cm.

FIG. 17 Illustration of the field and ion energy budget for null-helicity reconnection. Ions gained 65% of the dissipated field energy, and 48% of the dissipated field energy was converted non-classically.

FIG. 18 Increase in ion thermal energy normalized by reconnected field energy for varying resistivity enhancement factors (null-helicity), showing a correlation between the two.
FIGURES

(a) vacuum vessel

(b) flux core

FIG. 1.

shot 12249
$T_i = 12.6 \pm 1.1 \text{ eV}$
peak at $438.24 \pm 0.15$
$\chi = 1.02$
IDSP $R = 35-40 \text{ cm}$
IDSP $Z = -2.5-2.5 \text{ cm}$
t $= 260-270 \mu s$

FIG. 2.
FIG. 3.
FIG. 4.
FIG. 5.
FIG. 8.

FIG. 9.
Reconnected field energy

Ion flow energy

Electrons

Classical Ion Processes

Non-Classical Ion Processes

<1%

48%

35%

viscosity

Ion Thermal Energy*

64%

*including heat losses

FIG. 15.

FIG. 16.

FIG. 17.
FIG. 18.